

Status of 3ν mass-mixing parameters, circa 2013

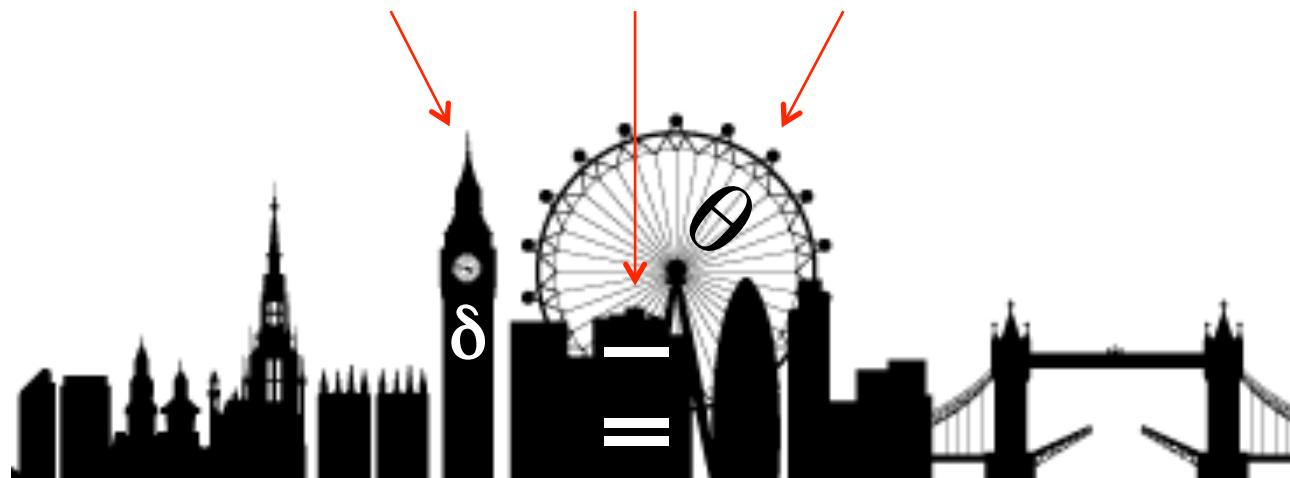


Elvio Lisi, INFN, Bari
Elvio Lisi, INFN, Bari

Outline:

- Intro: data, notation, methodology
- Global analysis: single parameters
- Global analysis: some covariances
- [Hierarchy issues at reactors] [if time allows it]
- Conclusions

Emphasis: CP phase, hierarchy, θ_{23} octant (unkowns)



Based on arXiv:13122878 [hep-ph], arXiv:13091638 [hep-ph]. Work done in collabor. with: F. Capozzi, G.L. Fogli, D. Montanino, A. Marrone, A. Palazzo

Data sets:

LBL Accelerators = **K2K + T2K + MINOS**

Solar = All Solar experiments

KL = KamLAND reactor expt

SBL Reactors = DChooz + RENO + DB

SK Atm = Super-K Atmospheric

3ν oscillation parameters: Notation

δm^2	=	Δm^2_{21}
$\theta_{12}, \theta_{23}, \theta_{13}, \delta$	=	as in PDB
δ range	=	$[0, 2\pi]$ (others prefer $[-\pi, +\pi]$)
Δm^2	=	$(\Delta m^2_{31} + \Delta m^2_{32})/2$

(All parameters free to float in the global fit)

Note: 1σ error on $\Delta m^2 \approx 0.07 \times 10^{-3} \text{ eV}^2 \approx \delta m^2$

Combined analysis of data sets: Methodology

LBL Accelerator data are dominantly sensitive to $(\Delta m^2, \theta_{23}, \theta_{13})$. But, accurate constraints on these parameters do need $(\delta m^2, \theta_{12})$ input from **Solar + KL** to compute sub-dominant effects.

Moreover: CP-violation is a genuine 3v effect, it would vanish in the approximation $\delta m^2 \sim 0$.

It makes sense to combine from the start:
LBL Acc + Solar + KL. Note: Solar + KL data carry a preference (“hint”) for $\sin^2 \theta_{13} \sim 0.02$

Combined analysis of data sets: Methodology

Analysis includes increasingly rich data sets:

LBL Acc + Solar + KL

LBL Acc + Solar + KL + SBL Reactor

LBL Acc + Solar + KL + SBL Reactor + SK Atm.

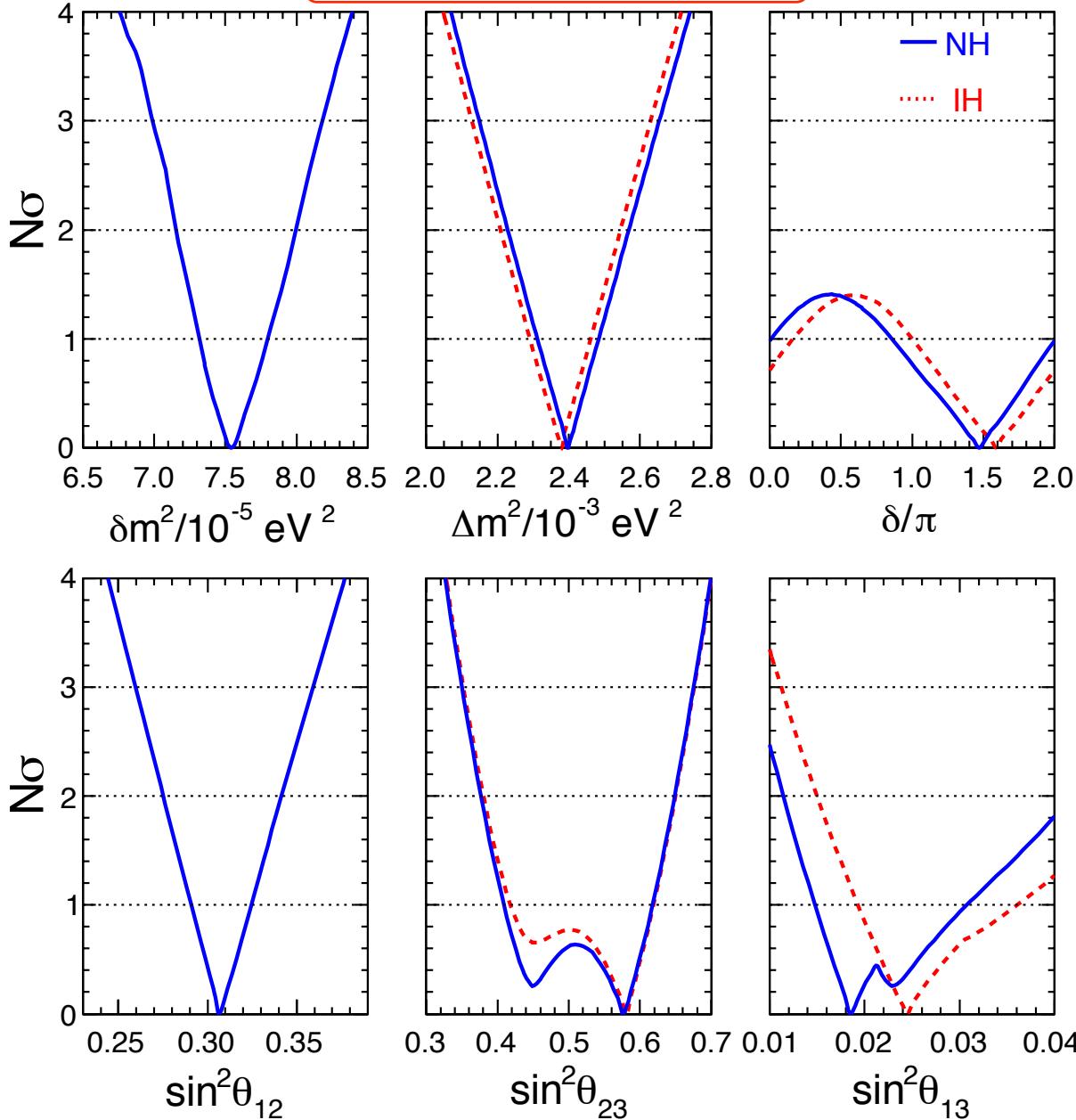
Figures: parameters not shown are marginalized away.

Contours are drawn at $\Delta\chi^2 = 1, 4, 9 \rightarrow$

$N\sigma = 1, 2, 3$ for projections over single parameters.

End of Intro. Results on single parameters →

LBL Acc + Solar + KL

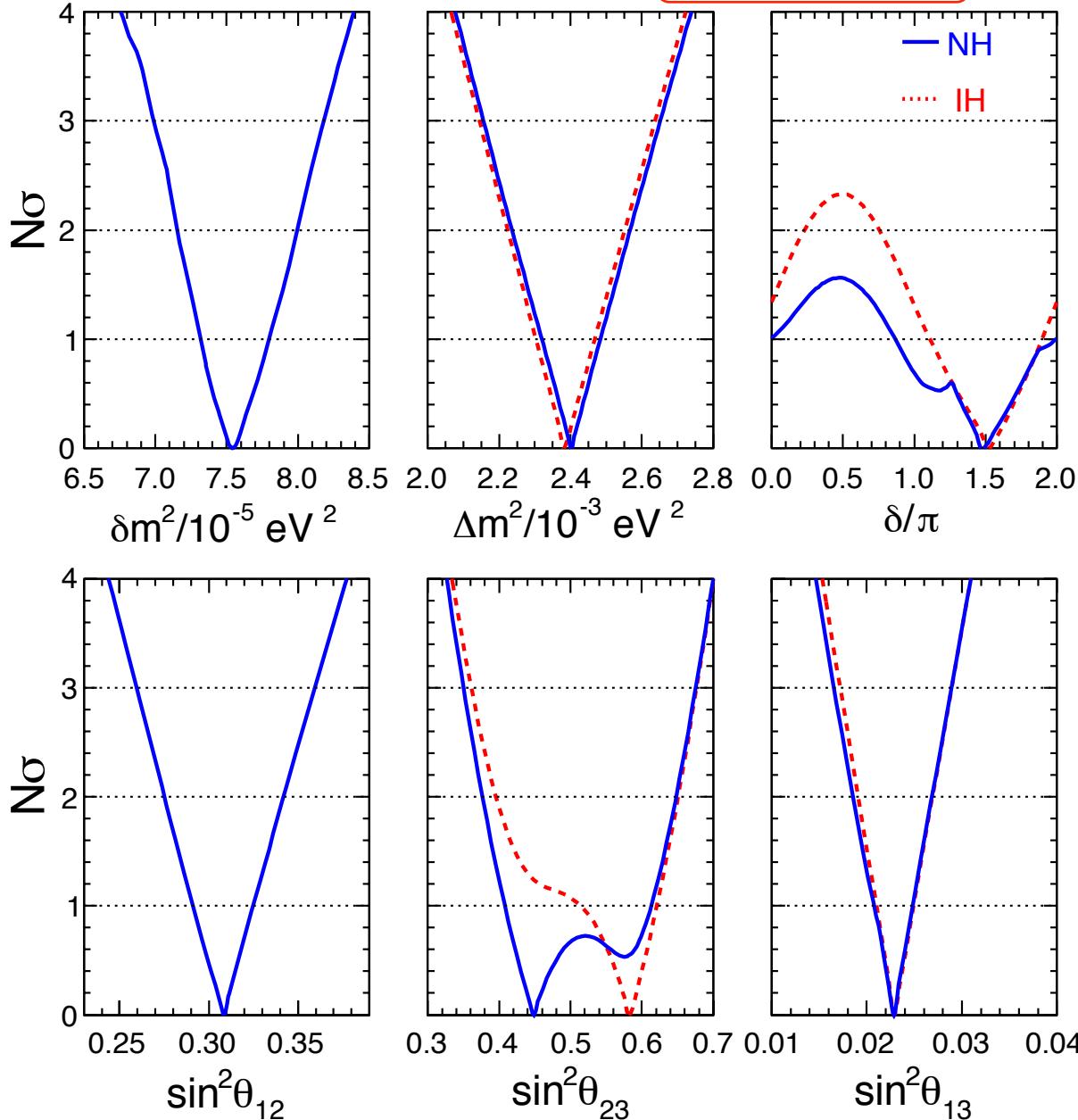


Upper and lower bound on all oscill. parameters but δ

Slight preference for $\delta \sim 1.5 \pi$

Slight preference for nonmaximal θ_{23} and for 2nd octant

LBL Acc + Solar + KL + SBL Reactors

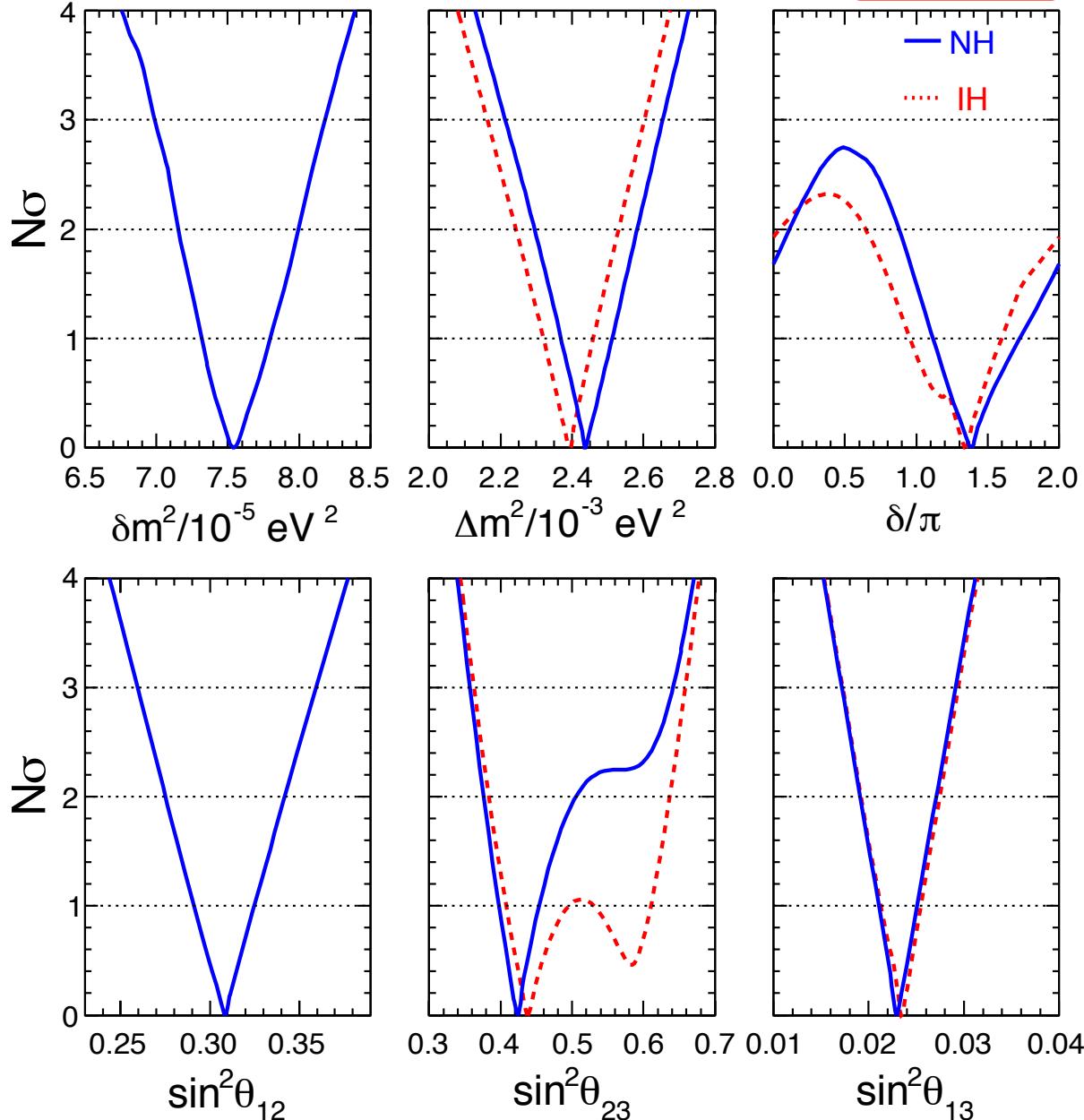


Strong, dominant
 θ_{13} bounds

Still a preference
for $\delta \sim 1.5 \pi$

Preference for
nonmaximal θ_{23}
but octant flips
with hierarchy

LBL Acc + Solar + KL + SBL Reactors + SK Atm



Some effects on the $\nu_\mu \rightarrow \nu_\tau$ dominant parameters ($\Delta m^2, \theta_{23}$)

Preference for $\delta \sim 1.4 \pi$ and, in NH, for $1 < \delta/\pi < 2$
 $(\sin \delta < 0 @ 90\% \text{ CL})$

Some preference for nonmaximal θ_{23} and for 1st octant, but weaker in IH

No ranges for single parameters (all data included):

[F. Capozzi, G.L. Fogli, E. Lisi, D. Montanino, A. Marrone, and A. Palazzo, arXiv:1312.2878]

TABLE I: Results of the global 3ν oscillation analysis, in terms of best-fit values and allowed 1, 2 and 3σ ranges for the 3ν mass-mixing parameters. See also Fig. 3 for a graphical representation of the results. We remind that Δm^2 is defined herein as $m_3^2 - (m_1^2 + m_2^2)/2$, with $+\Delta m^2$ for NH and $-\Delta m^2$ for IH. The CP violating phase is taken in the (cyclic) interval $\delta/\pi \in [0, 2]$. The overall χ^2 difference between IH and NH is insignificant ($\Delta\chi^2_{\text{I-N}} = +0.3$).

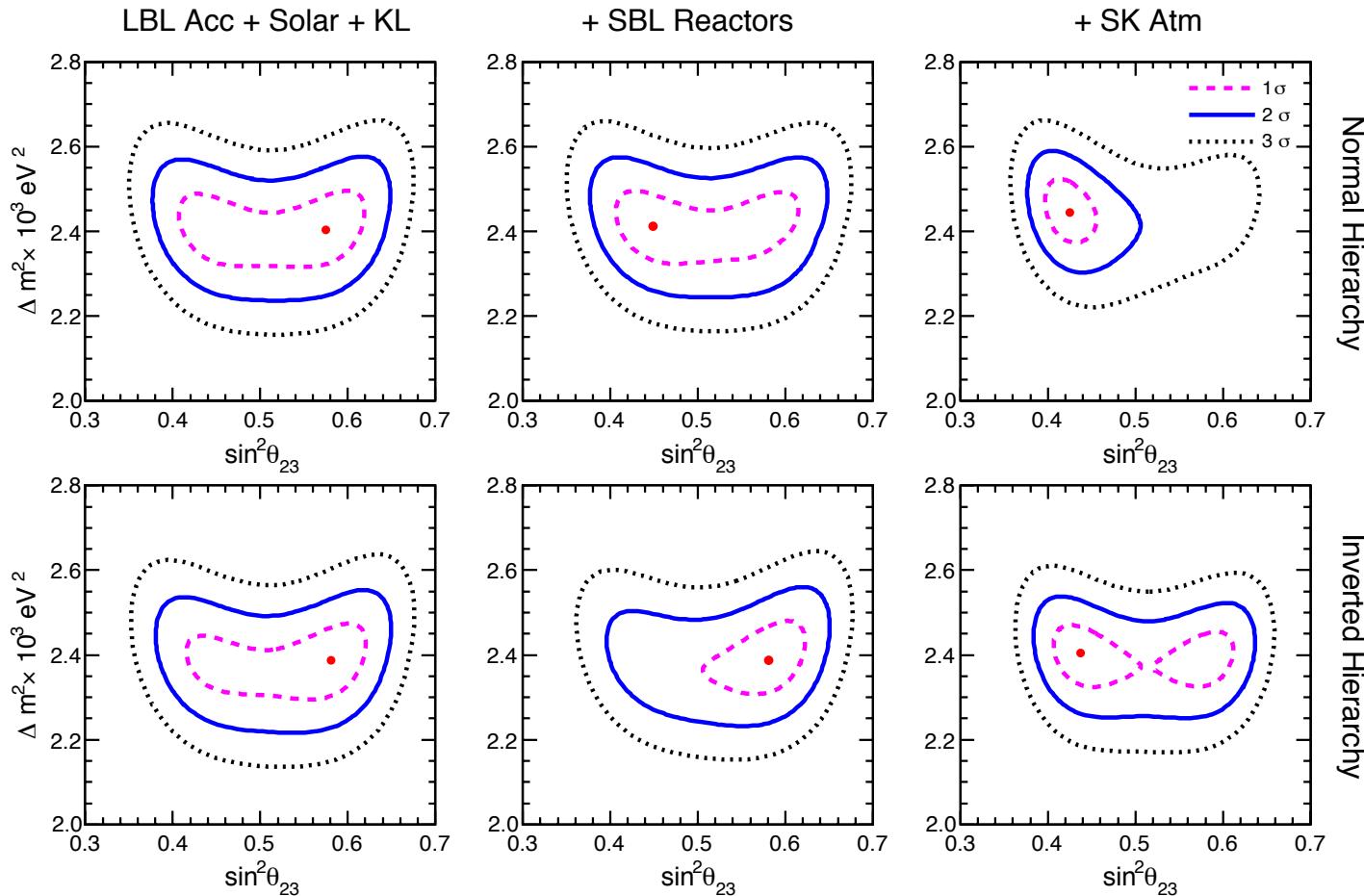
Parameter	Best fit	1σ range	2σ range	3σ range
$\delta m^2/10^{-5} \text{ eV}^2$ (NH or IH)	7.54	7.32 – 7.80	7.15 – 8.00	6.99 – 8.18
$\sin^2 \theta_{12}/10^{-1}$ (NH or IH)	3.08	2.91 – 3.25	2.75 – 3.42	2.59 – 3.59
$\Delta m^2/10^{-3} \text{ eV}^2$ (NH)	2.44	2.38 – 2.52	2.30 – 2.59	2.22 – 2.66
$\Delta m^2/10^{-3} \text{ eV}^2$ (IH)	2.40	2.33 – 2.47	2.25 – 2.54	2.17 – 2.61
$\sin^2 \theta_{13}/10^{-2}$ (NH)	2.34	2.16 – 2.56	1.97 – 2.76	1.77 – 2.97
$\sin^2 \theta_{13}/10^{-2}$ (IH)	2.39	2.18 – 2.60	1.98 – 2.80	1.78 – 3.00
$\sin^2 \theta_{23}/10^{-1}$ (NH)	4.25	3.98 – 4.54	3.76 – 5.06	3.57 – 6.41
$\sin^2 \theta_{23}/10^{-1}$ (IH)	4.37	4.08 – 4.96 \oplus 5.31 – 6.10	3.84 – 6.37	3.63 – 6.59
δ/π (NH)	1.39	1.12 – 1.72	0.00 – 0.11 \oplus 0.88 – 2.00	—
δ/π (IH)	1.35	0.96 – 1.59	0.00 – 0.04 \oplus 0.65 – 2.00	—

Fractional uncertainties (defined as 1/6 of 3σ ranges):

δm^2	2.6 %
Δm^2	3.0 %
$\sin^2 \theta_{12}$	5.4 %
$\sin^2 \theta_{13}$	8.5 %
$\sin^2 \theta_{23}$	~11 %

Selected parameter covariances →

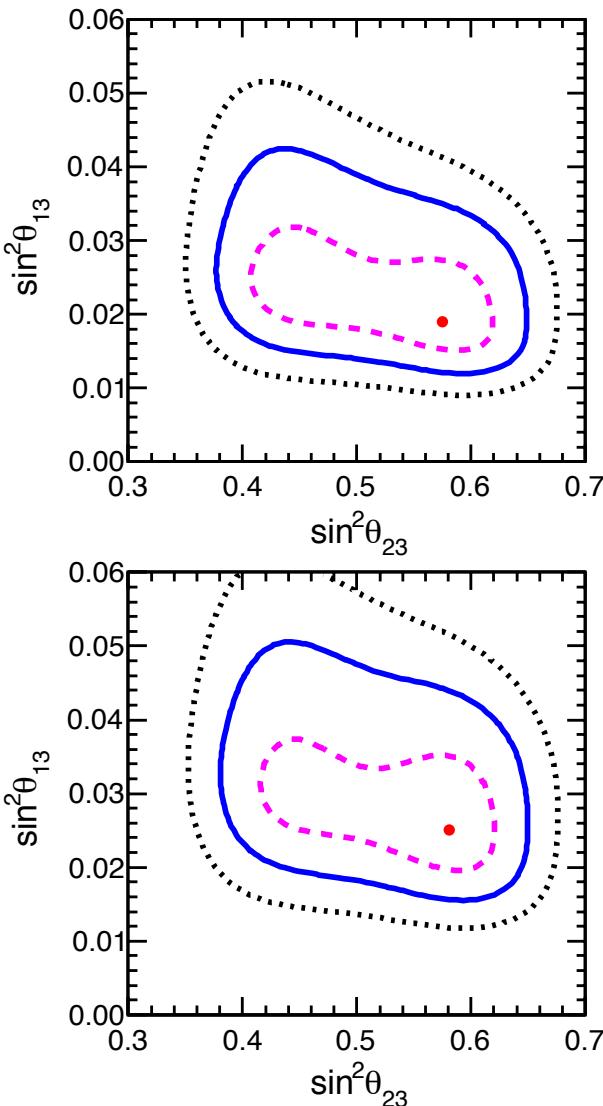
The θ_{23} octant “flip” in a more familiar plane:



but ... easier to understand in $(\theta_{23}, \theta_{13})$ plane

[Note, however, relevance of future reactor data in breaking correlations in above figure.]

LBL Acc + Solar + KL



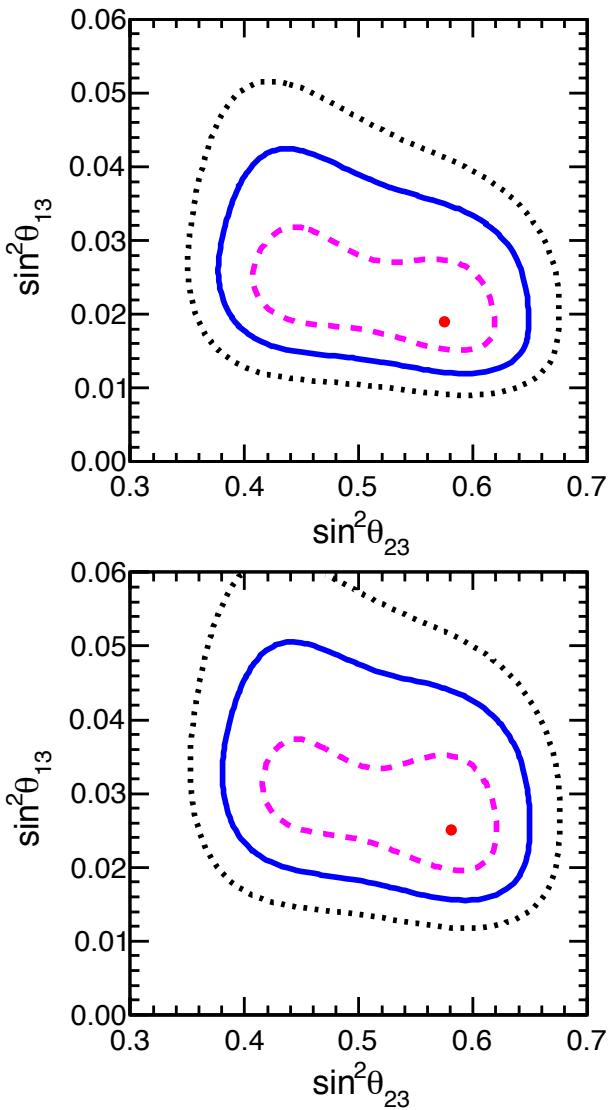
MINOS disappearance prefers nonmaximal mixing (and wins over T2K preference for \sim maximal) \rightarrow two degenerate minima for θ_{23}

T2K + MINOS appearance anticorrelate the minima with θ_{13} : the higher θ_{23} , the lower θ_{13}
[appearance amplitude $\sim \sin^2\theta_{23}\sin^2(2\theta_{13})$]

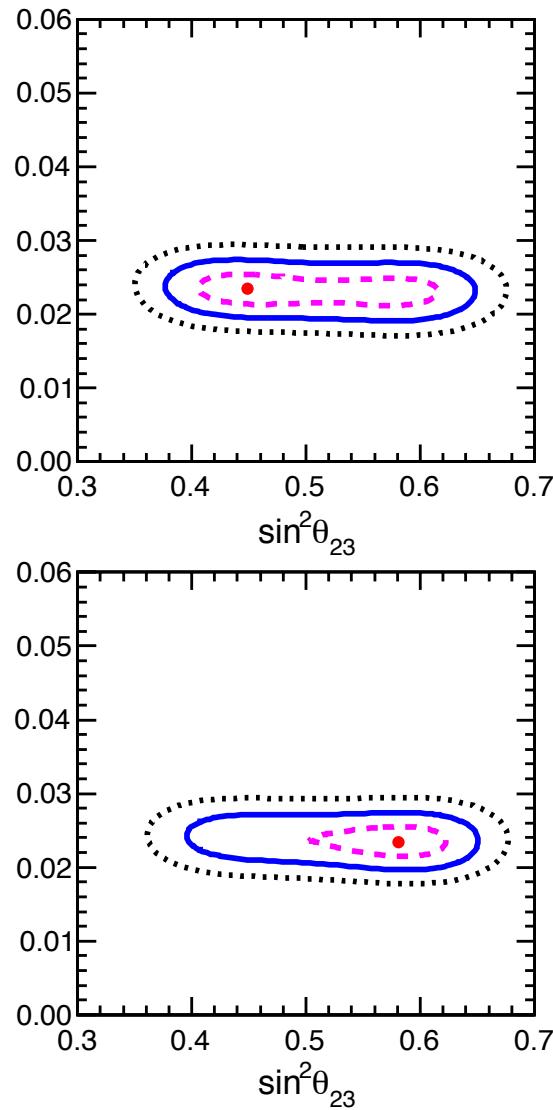
Contours extend to relatively high $\sin^2\theta_{13}$ to accommodate the relatively “strong” T2K appearance signal, especially in IH

In the combination, Solar + KL data lift the degeneracy and prefer the second octant solution, associated with “low” $\sin^2\theta_{13} \sim 0.02$

LBL Acc + Solar + KL

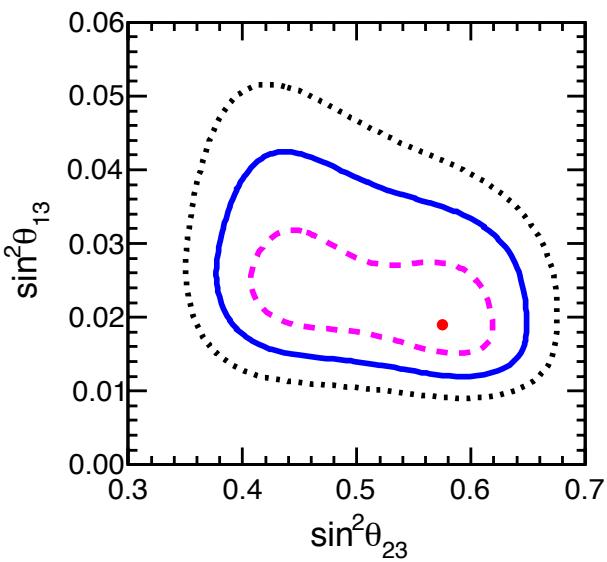


+ SBL Reactors

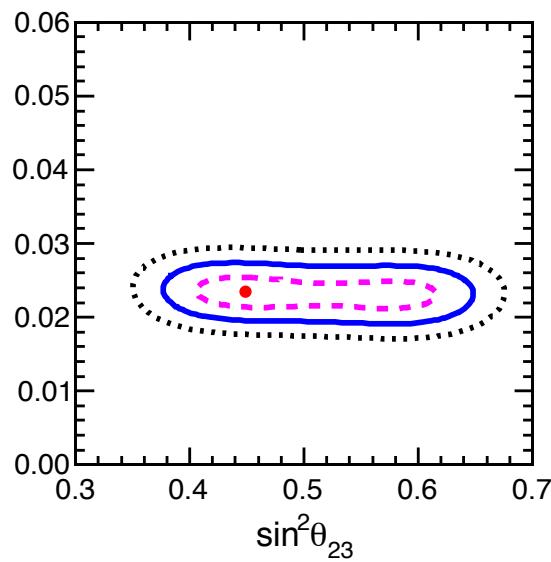


Reactor data prefer $\sin^2\theta_{13} \sim 0.023$, slightly higher than Solar+KL: enough to flip the octant in NH, but not enough to do so in IH.

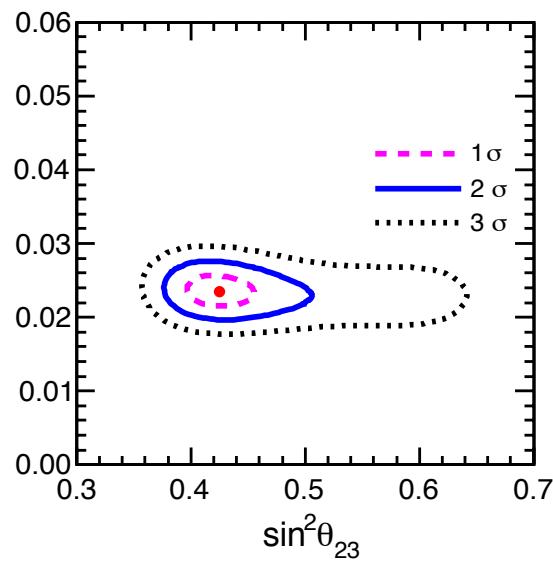
LBL Acc + Solar + KL



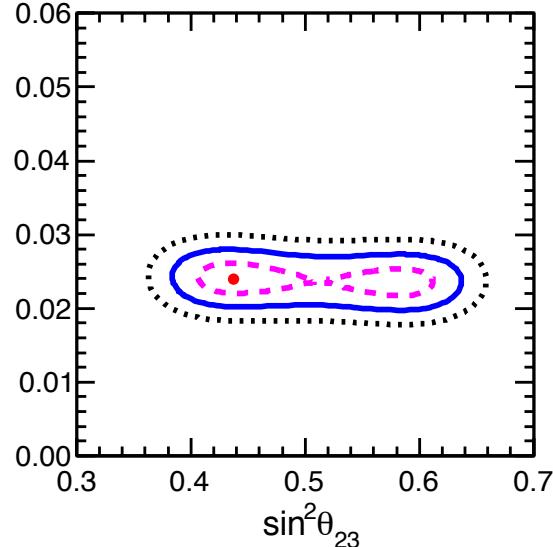
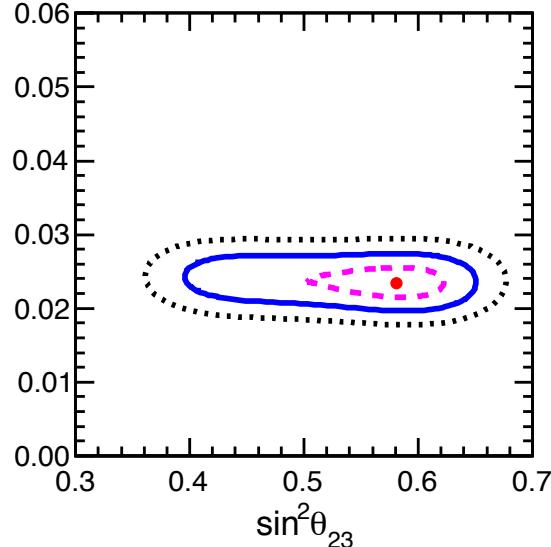
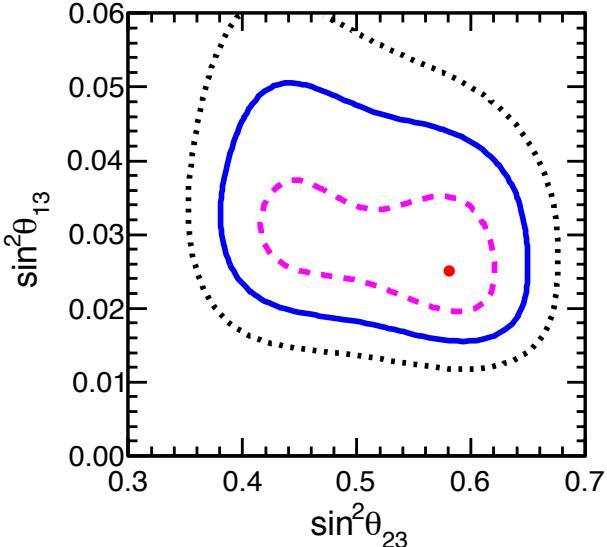
+ SBL Reactors



+ SK Atm



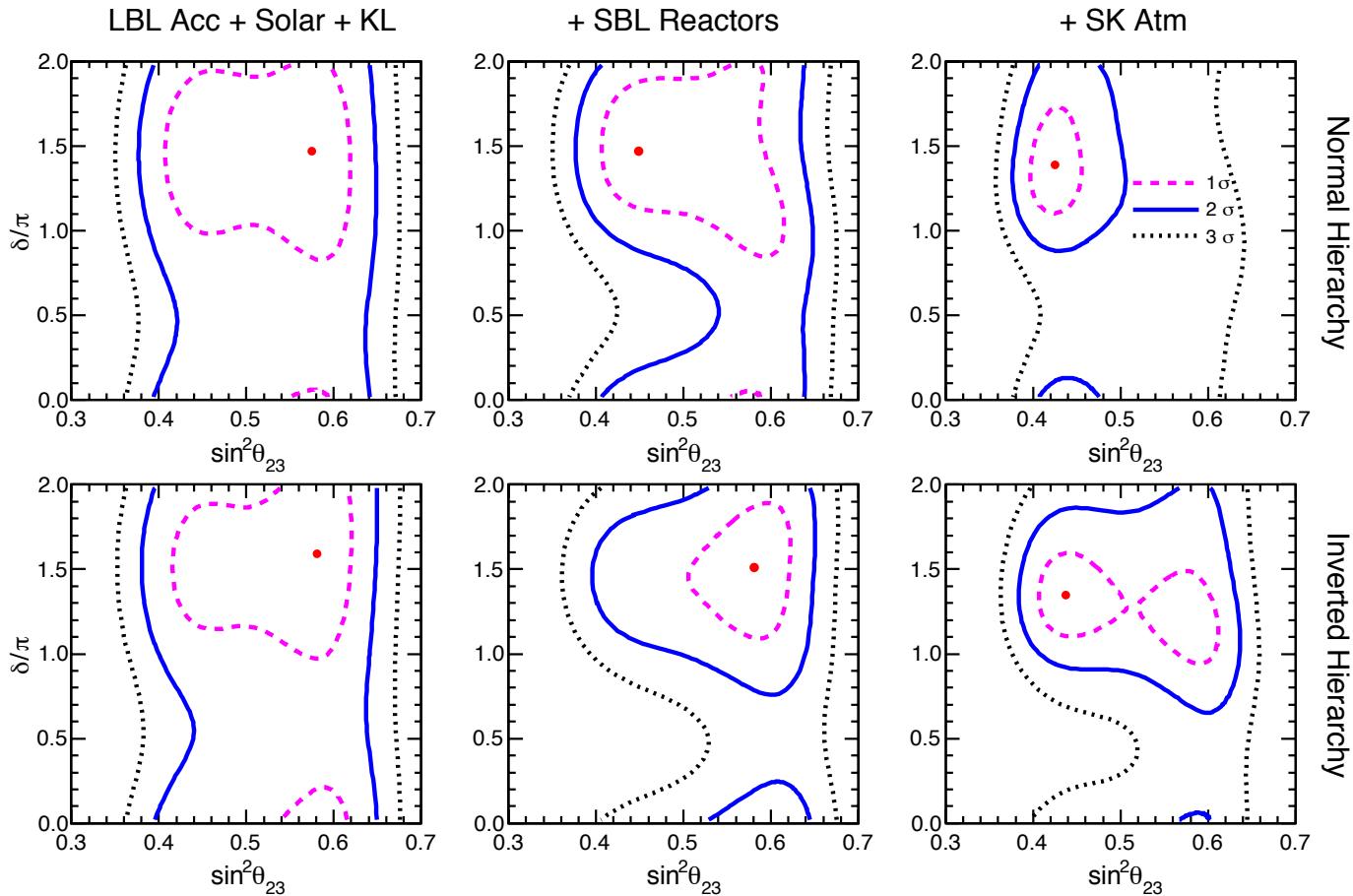
Normal Hierarchy



Inverted Hierarchy

SK atm: We continue to find an overall preference of atmospheric data for the first octant – which currently wins over other data.

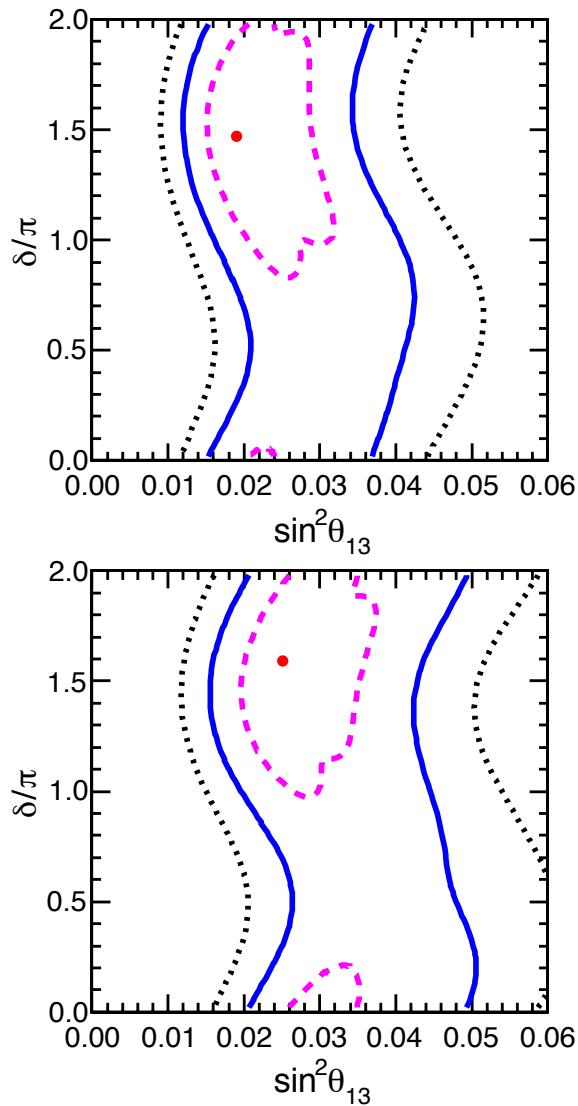
Interpretation of $\delta \sim 1.4\pi$ preference ...



... easier by looking at (δ, θ_{13}) correlations

[Note, however, strong asymmetry of allowed regions with respect to octant.]

LBL Acc + Solar + KL



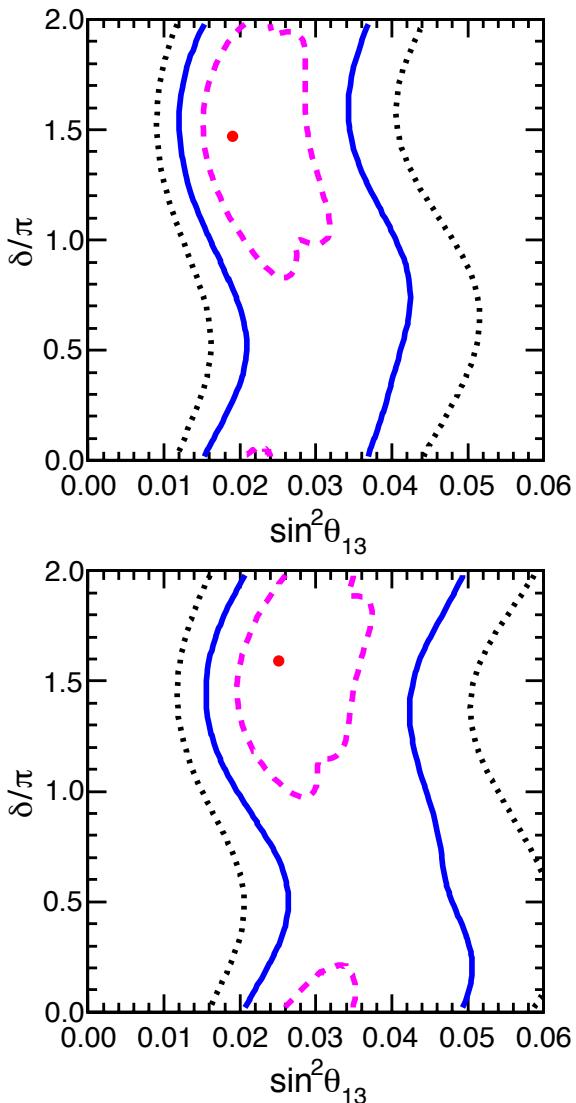
Each wavy band is in part determined by superposition of “two bands” for the two θ_{23} octants

For the relatively “low” value $\sin^2\theta_{13} \sim 0.02$ preferred by Solar + KL data, appearance ν signal in T2K maximized by subleading CP-odd term for $\sin\delta < 0$ [i.e., $1 < \delta/\pi < 2$]

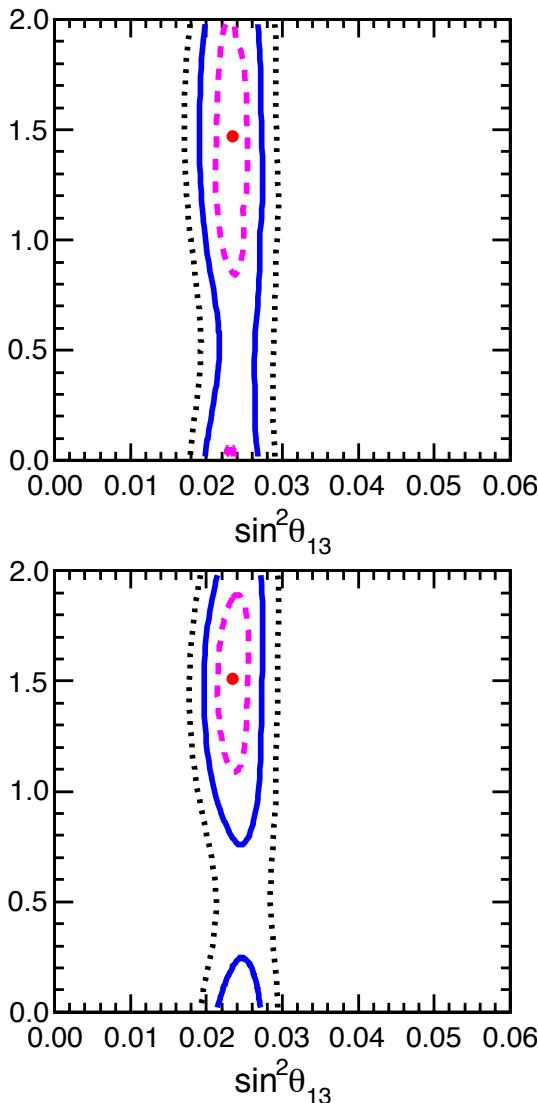
Best agreement with relatively “strong” T2K appearance signal is for $\delta/\pi \sim 1.5$, irrespective of the hierarchy.

This trend wins over weaker MINOS appearance signal, which generally prefers $\sin\delta < 0$ at best fit.

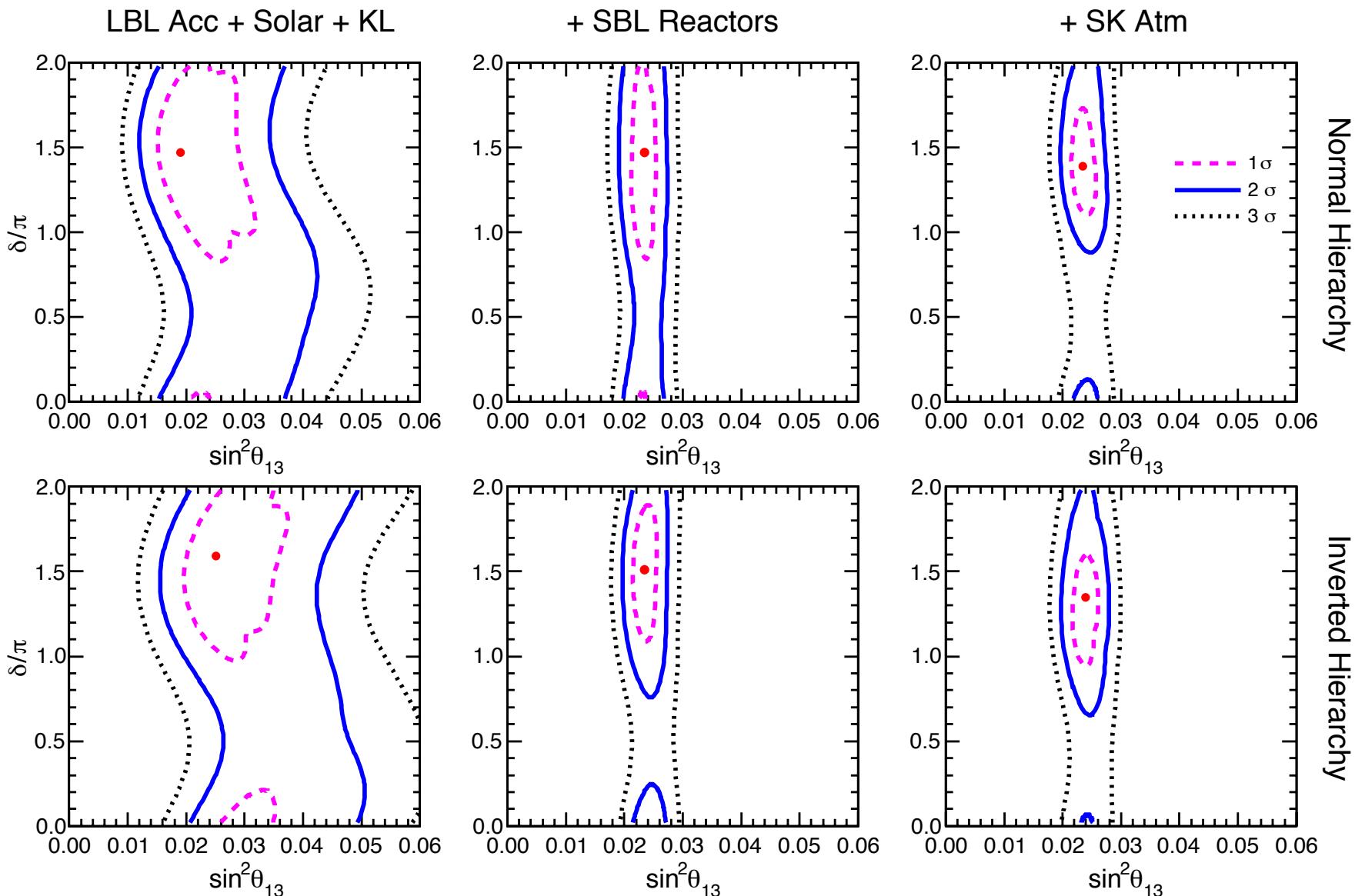
LBL Acc + Solar + KL



+ SBL Reactors



**Reactor data shrink
the band around
 $\sin^2\theta_{13} \sim 0.023$,
a bit higher than
Solar+SK but still
on the leftmost
side of the band:
preference for
 $\delta/\pi \sim 1.5$ persists**



SK atm: We continue to find an overall preference for $\delta/\pi \sim 1$ (with $\delta \sim 0$ disfavored). In combination, $\delta/\pi \sim 1.4$ and $\sin \delta < 0$ favored.

What about NH vs IH?

Figure of merit: $\Delta\chi^2 = \chi^2_{\min}(\text{NH}) - \chi^2_{\min}(\text{IH})$

LBL Acc + Solar + KL	:	$\Delta\chi^2 = +1.3$
LBL Acc + Solar + KL + SBL Reactor	:	$\Delta\chi^2 = +1.4$
LBL Acc + Solar + KL + SBL Reactor + SK Atm.	:	$\Delta\chi^2 = -0.3$

No significant sensitivity yet.

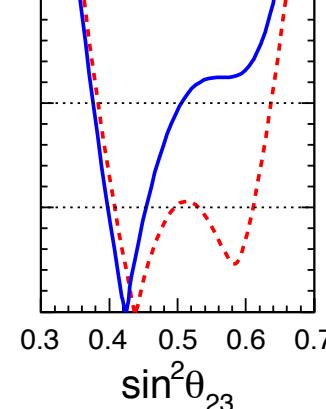
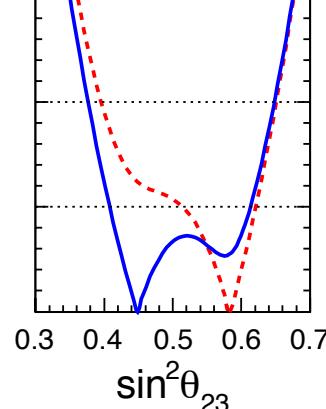
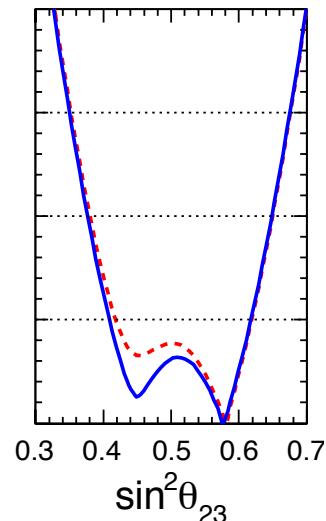
Unknown parameters: summary of our results →

LBL + Sol. + KL

+ SBL Reactor

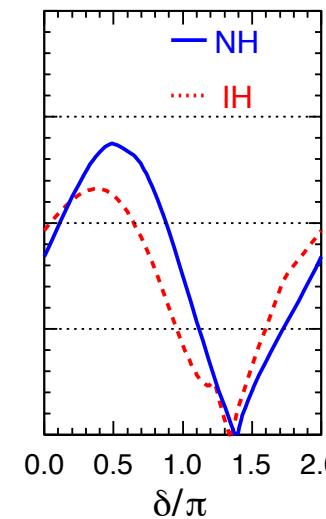
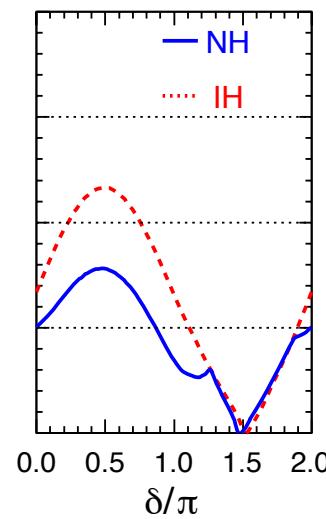
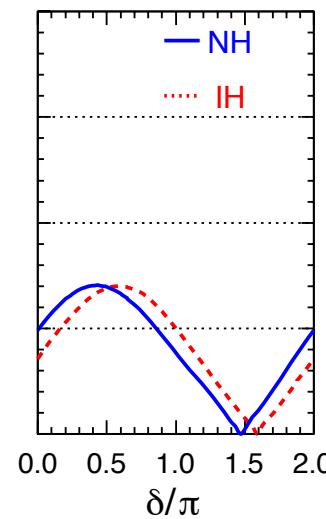
+ SK Atmos.

θ_{23}



Apparent instability on octant

δ



Weak but intriguing synergy on CP phase

$\Delta\chi^2$
(IH-NH)

-1.3



-1.4

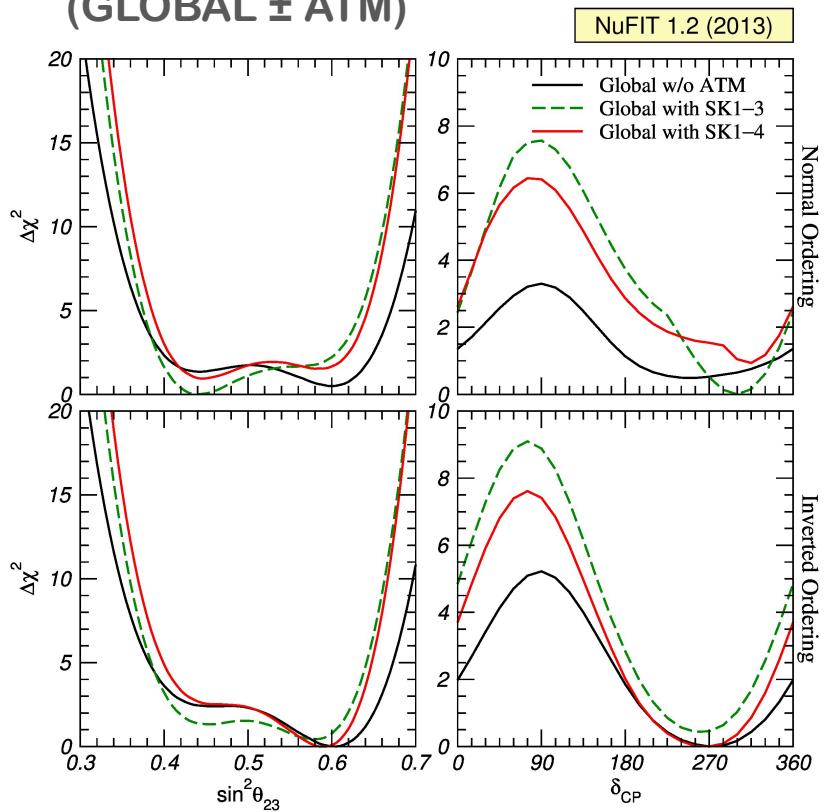


+0.3

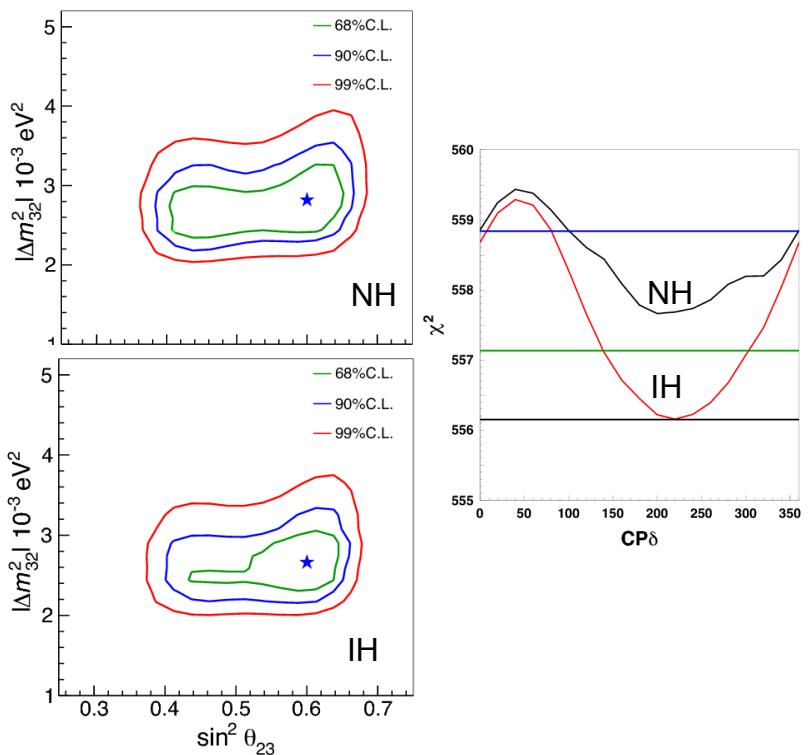
No hint NH/IH

Unknown parameters: comparison wrt ...

Gonzalez-Garcia et al. 2013
(GLOBAL \pm ATM)

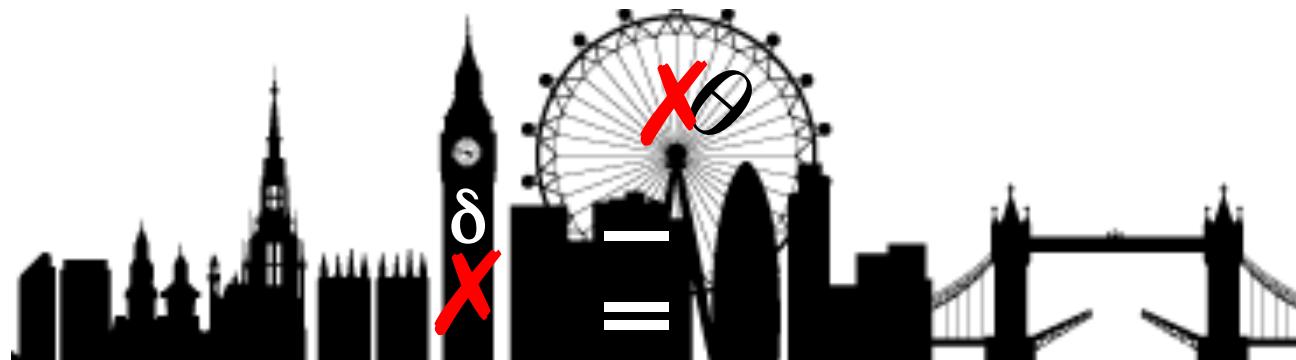
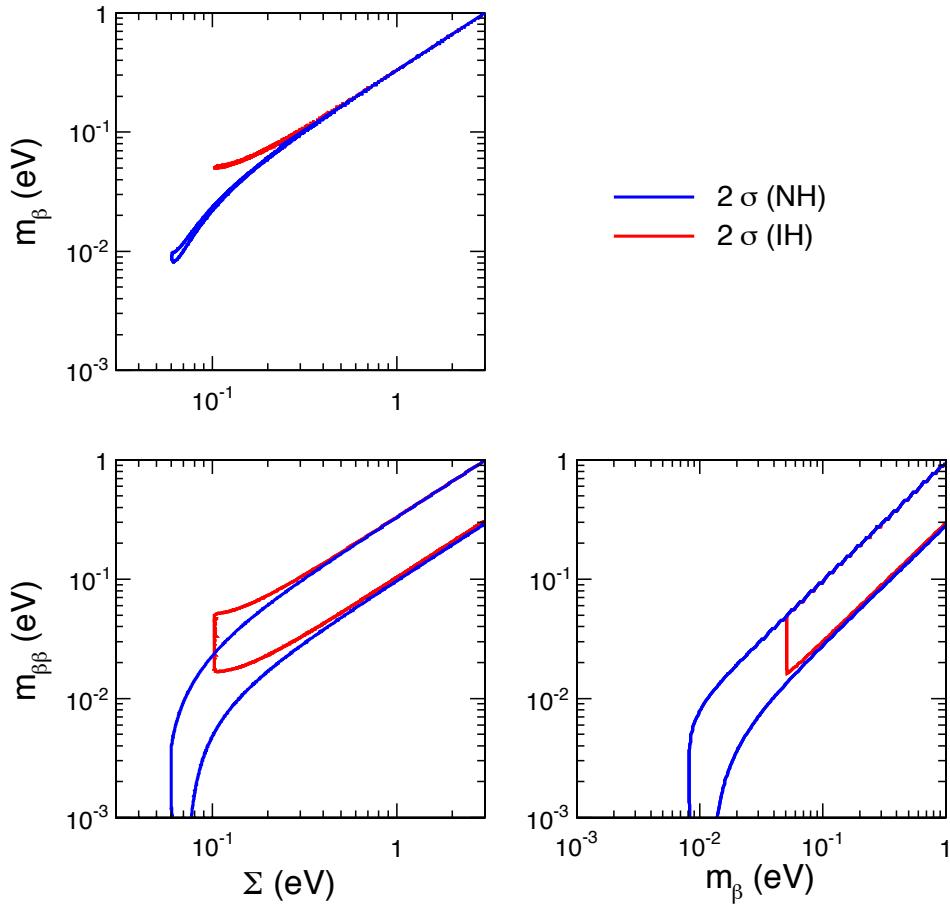
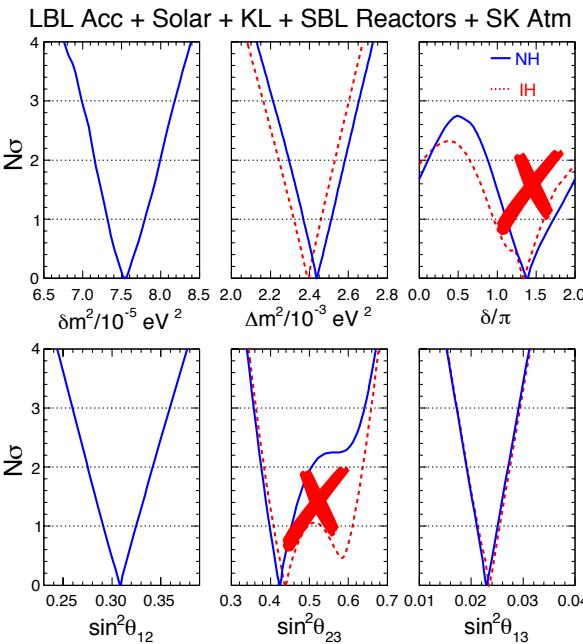


Super-Kamiokande, arXiv:1310.6677
(REACTOR + ATM)

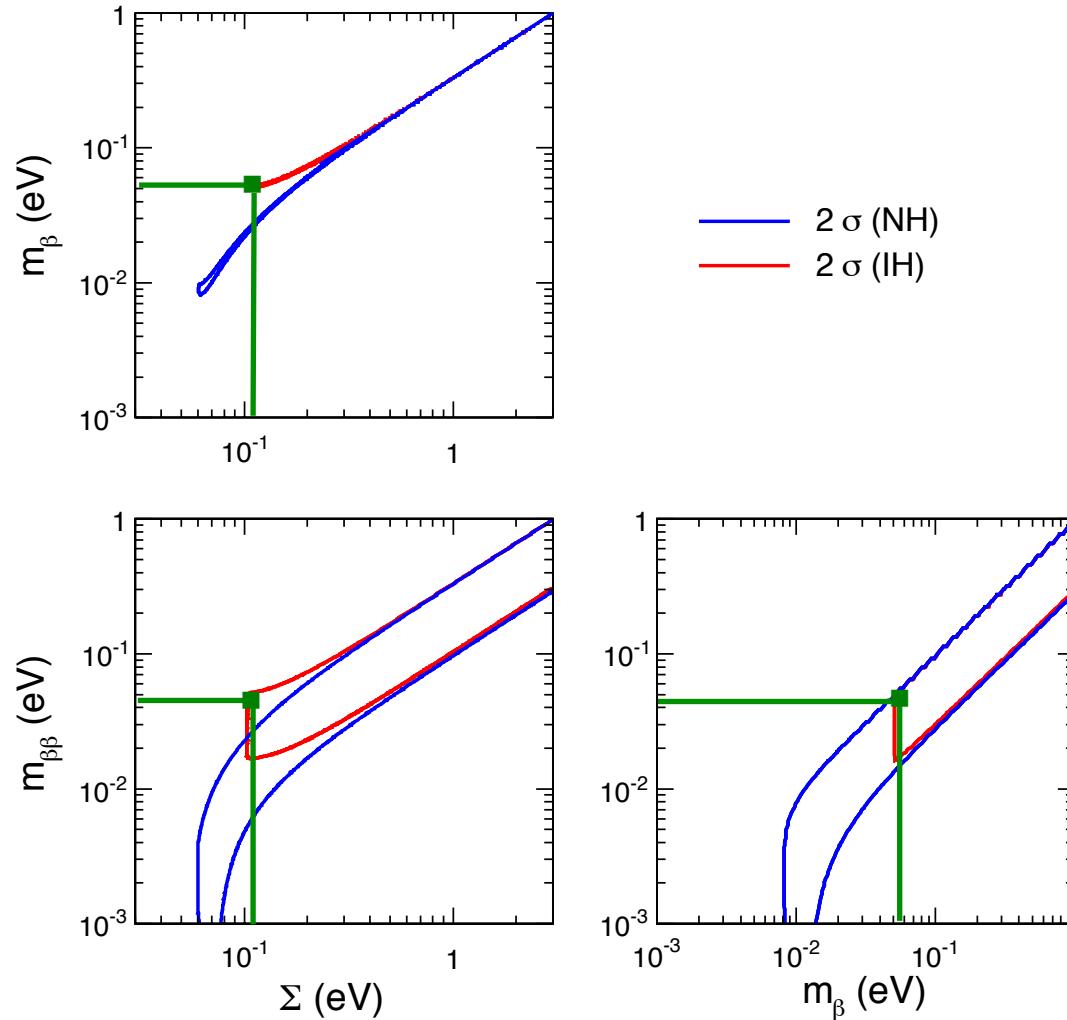


Relative fluctuations of best-fit octant and hierarchy within errors,
but interesting convergence on best-fit CP phase at $\sin\delta < 0$

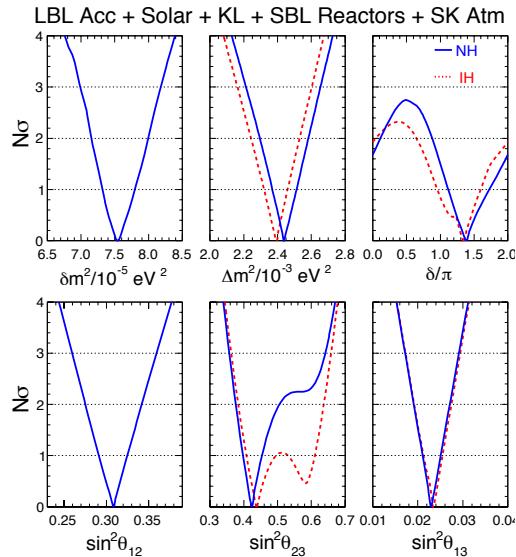
Implications for absol. ν masses:



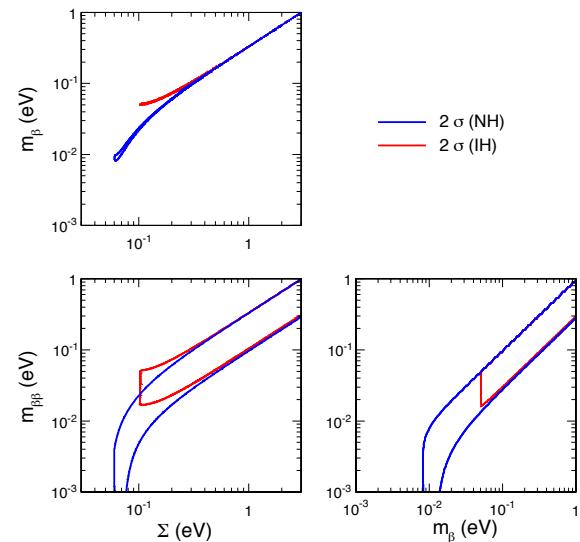
Absolute ν mass observables may provide an independent handle, or at least constraints, on NH vs IH discrimination. An ideal case:



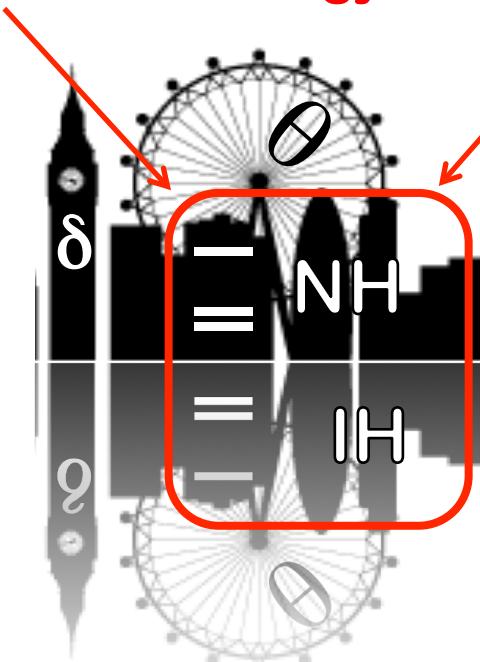
Oscillation data



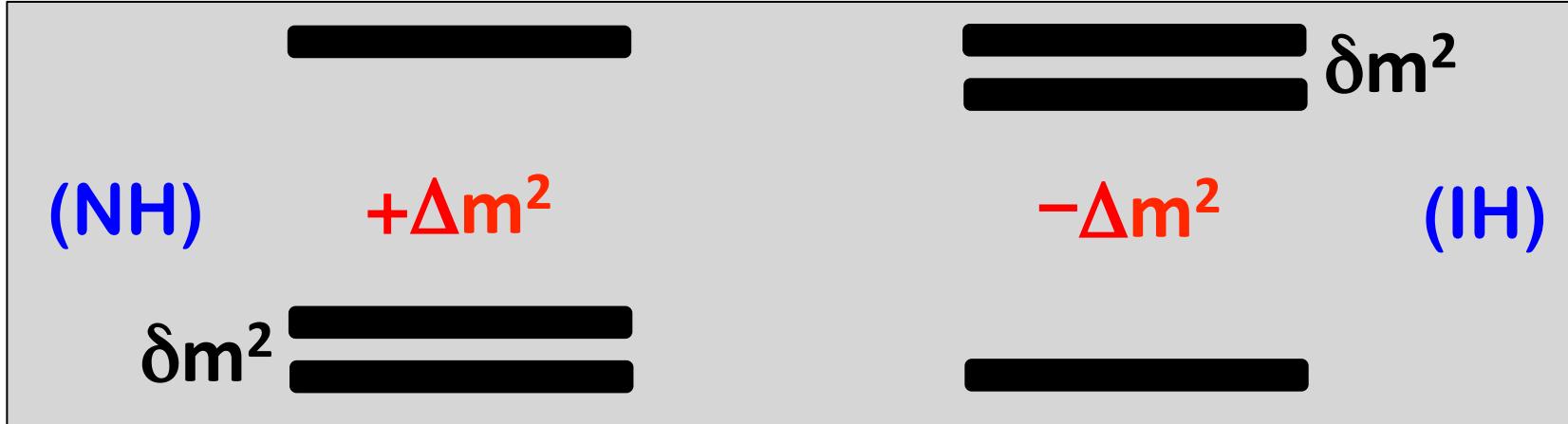
Nonoscillation data



Sinergy



Neutrino flavor oscillations can probe the hierarchy...



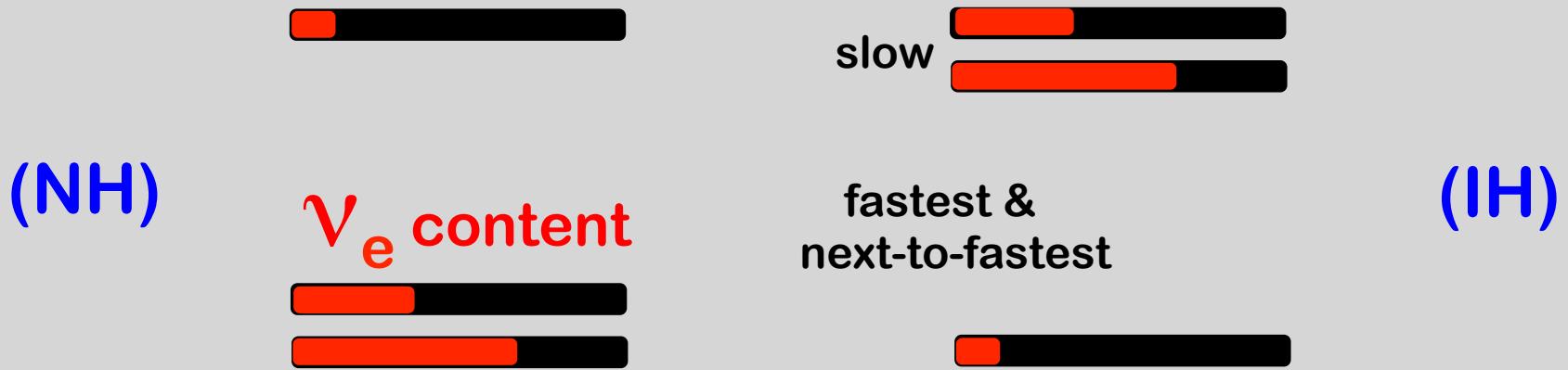
... if oscillations driven by $\pm \Delta m^2$ interfere with oscillations driven by another “squared mass gap” Q with known sign. Three options:

$$Q = \delta m^2 \quad (\text{focus of the following slides} \rightarrow)$$

$$Q = 2\sqrt{2} G_F N_e E \quad (\text{matter effects in Earth or SNe})$$

$$Q = 2\sqrt{2} G_F N_\nu E \quad (\text{collective effects in SNe})$$

Simple physics: One slow & two fast oscillations



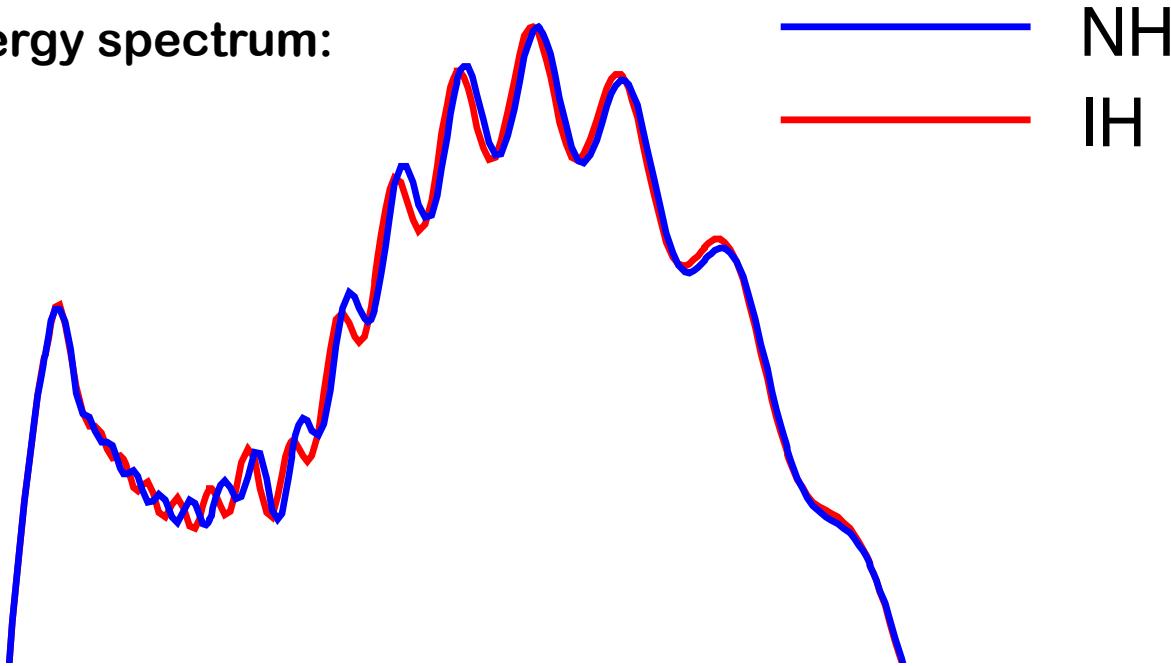
ν_e oscillation amplitude = product of two red bars

Under hierarchy swap:

Amplitude of the slow oscillation does not change, while the (different) amplitudes of the fastest and next-to-fastest oscillations are interchanged

(unless closest red bars were equal, i.e., $\theta_{12}=\pi/4$)

Typical energy spectrum:



Peaks in oscillated reactor spectrum at $O(50)$ km are displaced from **NH** to **IH**: one cannot exactly mimic the other in the ideal case of no errors, even if all other osc. parameters are varied.

However, within statistical errors, a “morphing” **NH \leftrightarrow IH** may be tolerated to some extent: **debate on statistical significance of $\Delta\chi^2$ difference**.

We have discussed* an analytical form for the reactor 3ν oscill. probability, including matter & damping effects, which allows a **continuous morphing between NH and IH**:

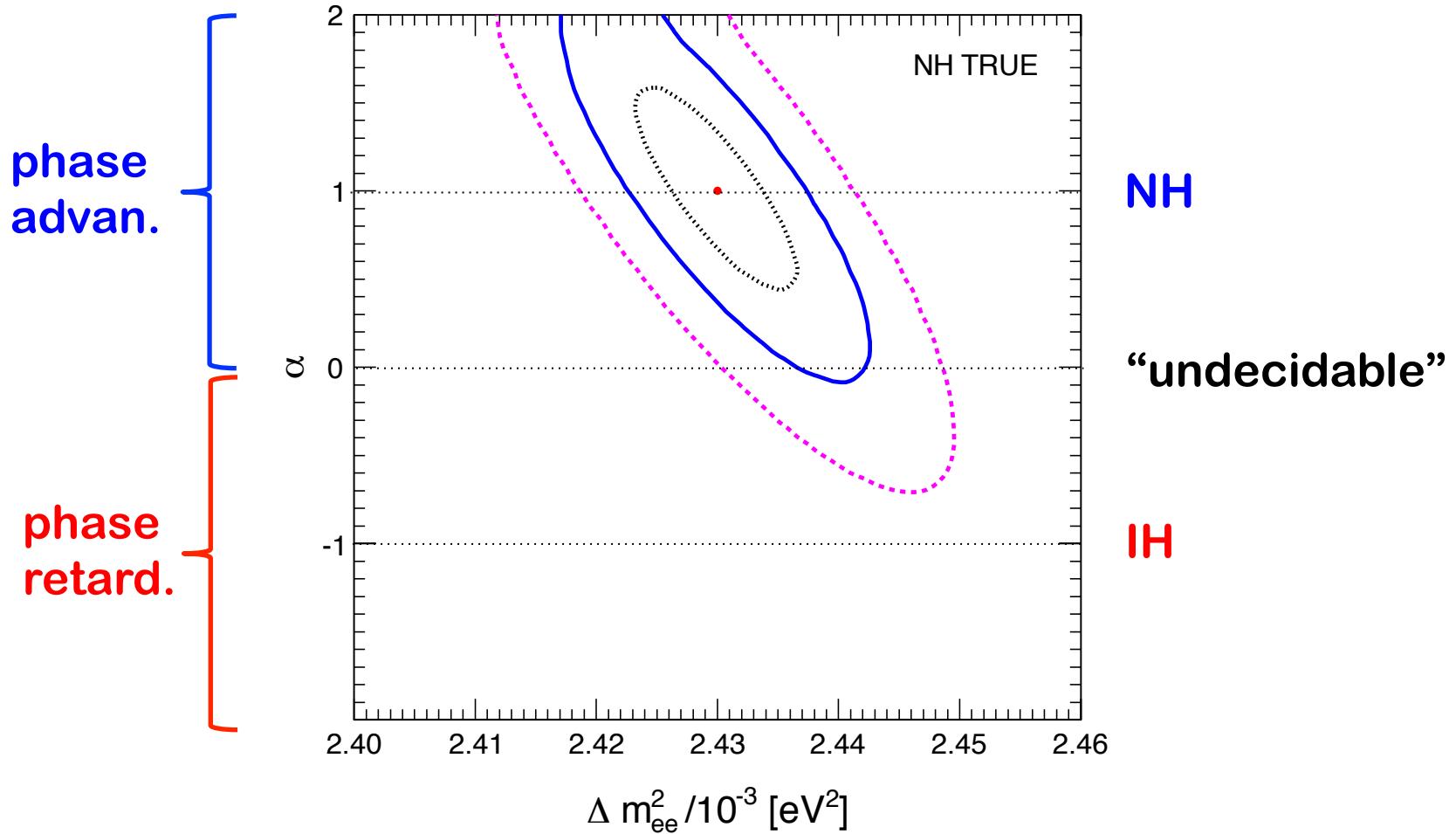
$$P_{\text{mat}}^{3\nu} \simeq c_{13}^4 P_{\text{mat}}^{2\nu} + s_{13}^4 + 2s_{13}^2 c_{13}^2 \sqrt{P_{\text{mat}}^{2\nu}} w \cos(2\Delta_{ee} + \cancel{\alpha}\varphi)$$

$$\alpha = \pm 1 \text{ (NH/IH)} \longrightarrow \alpha = \text{free}$$

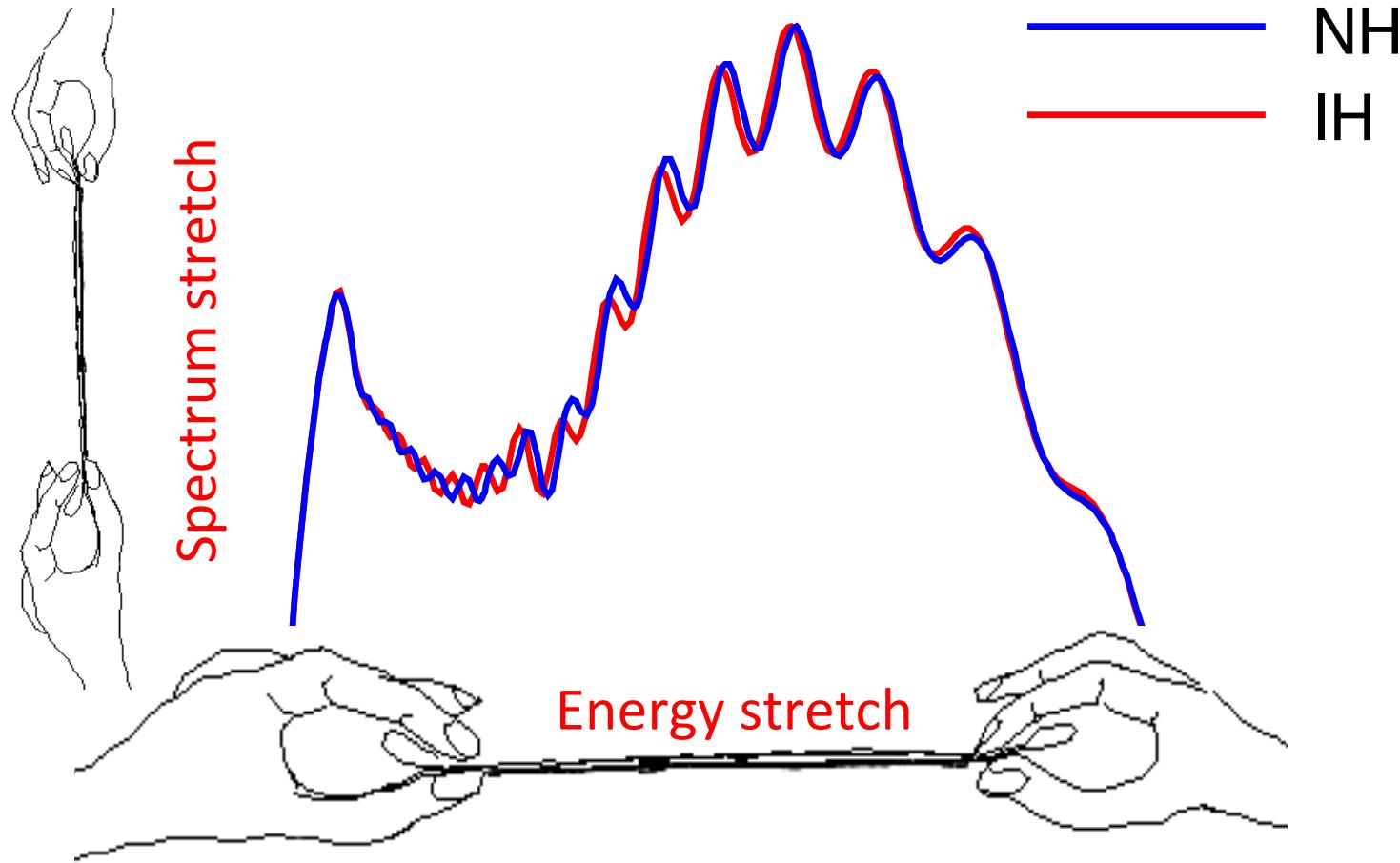
Prospective results for a JUNO-like setting (5 yrs) →
[including stat. errors + osc. par. errors + normalization syst.]

* F. Capozzi, E. Lisi, A. Marrone, arXiv:1309.1638, to appear in PRD (2013)

$(\Delta m^2_{ee}, \alpha)$ fit for true NH



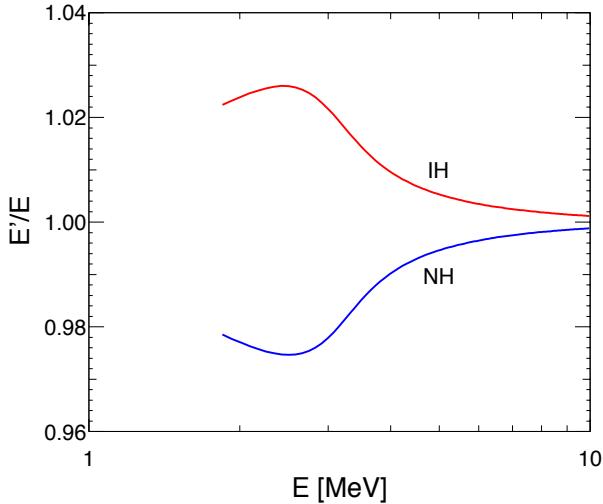
Although the “wrong” hierarchy is $\sim 3.4\sigma$ away from the true one, the experiment is already compromised when the “undecidable” case is reached at $\sim 1.7\sigma = \text{effective sensitivity}$. Under debate!



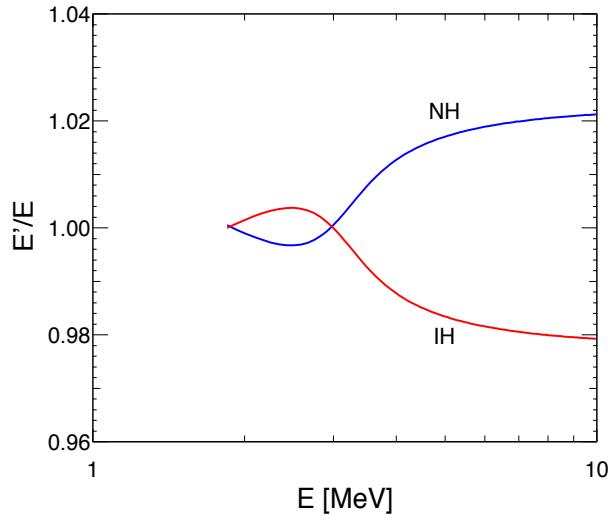
In the previous fit we have not included possible nonlinear stretch of the x-axis (**energy scale systematics**) or of the y axis (**spectrum shape systematics**) which could allow an easier “**morphing**” NH/IH by realigning the peaks and renormalizing their heights.

There are infinite ways to do so! Just two examples, involving x,y stretch localized at low E or at high E:

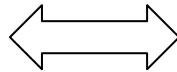
low E



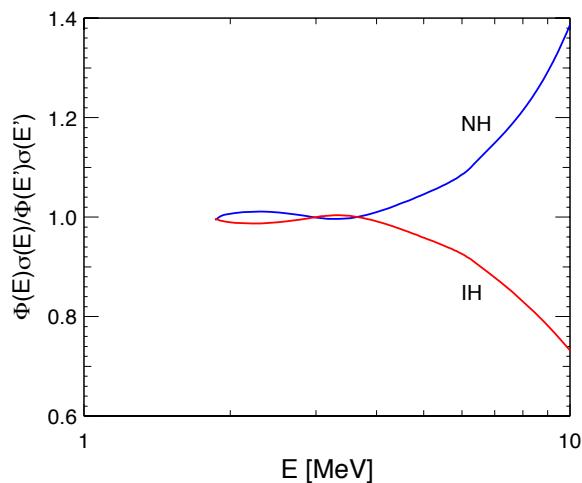
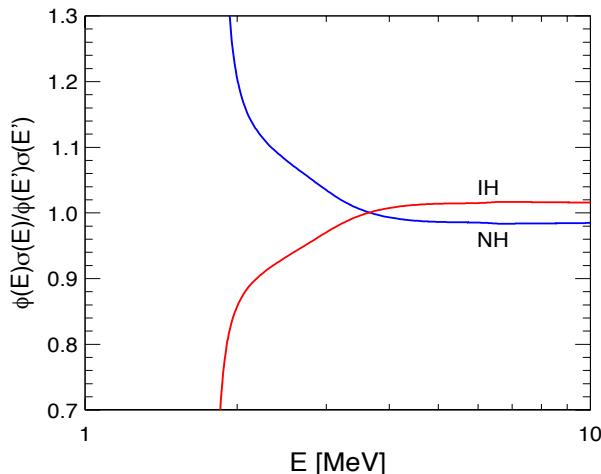
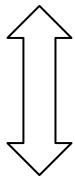
high E



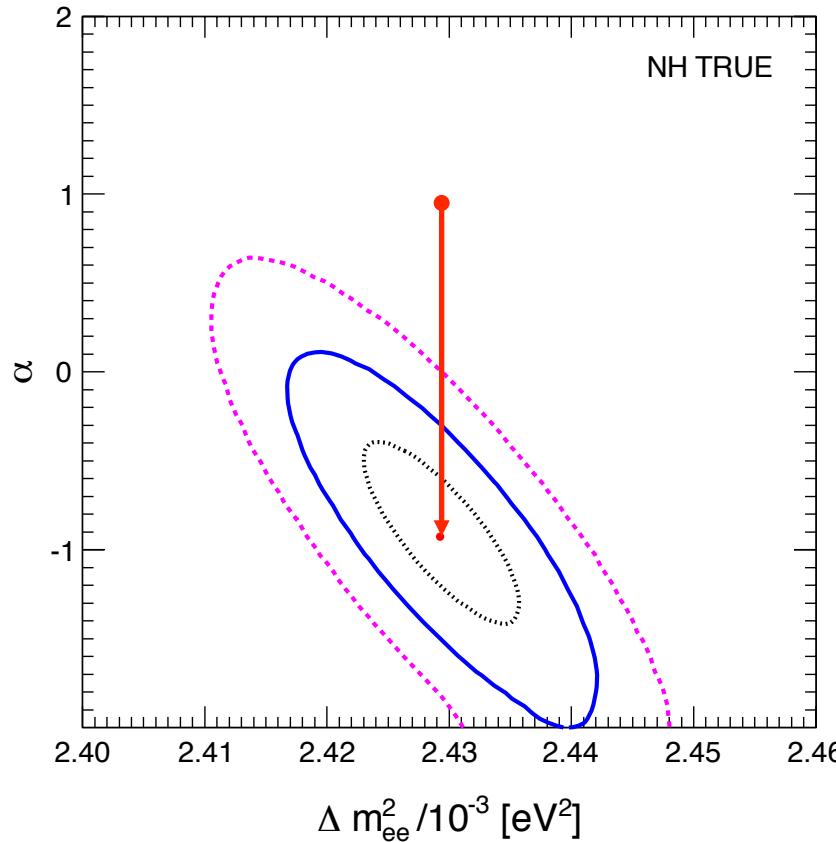
X-axis
stretch



Y-axis
stretch



Stretch effect: true hierarchy \rightarrow wrong hierarchy



Quantifying these **systematics** and their possible functional shapes represent a new frontier of precision reactor experiments and of their analysis.

Conclusions:

Beautiful neutrino experiments have sketched the current “**three-neutrino skyline**”, sometimes with amazing accuracy. But basic pieces, which may profoundly change the landscape, are still missing. In this context, global analyses may provide some guidance on emerging profiles and on unprecedented challenges in the field. In this process, let’s never forget that...



Conclusions:

Beautiful neutrino experiments have sketched the current “**three-neutrino skyline**”, sometimes with amazing accuracy. But basic pieces, which may profoundly change the landscape, are still missing. In this context, global analyses may provide some guidance on emerging profiles and on unprecedented challenges in the field. In this process, let’s never forget that ... **surprises may appear & may open new gates!**



Thank you for your attention.

Synopsis:

[F. Capozzi, G.L. Fogli, E. Lisi, D. Montanino, A. Marrone, and A. Palazzo, arXiv:1312.2878]

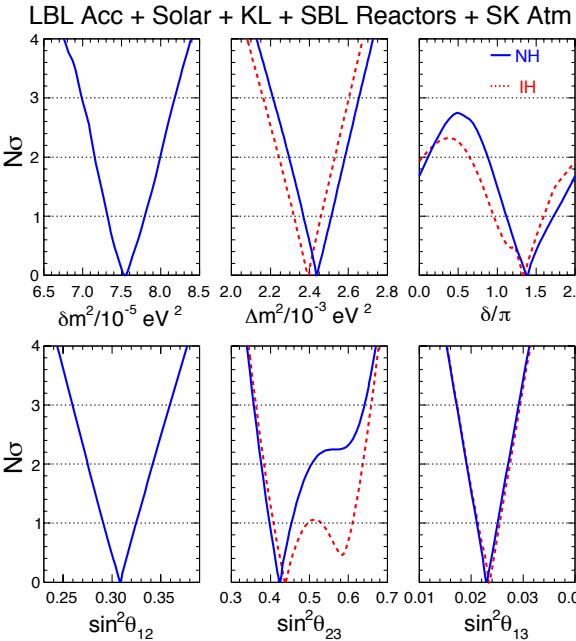


TABLE I: Results of the global 3ν oscillation analysis, in terms of best-fit values and allowed 1, 2 and 3σ ranges for the 3ν mass-mixing parameters. See also Fig. 3 for a graphical representation of the results. We remind that Δm^2 is defined herein as $m_3^2 - (m_1^2 + m_2^2)/2$, with $+\Delta m^2$ for NH and $-\Delta m^2$ for IH. The CP violating phase is taken in the (cyclic) interval $\delta/\pi \in [0, 2]$. The overall χ^2 difference between IH and NH is insignificant ($\Delta\chi^2_{I-N} = +0.3$).

Parameter	Best fit	1σ range	2σ range	3σ range
$\delta m^2/10^{-5} \text{ eV}^2$ (NH or IH)	7.54	7.32 – 7.80	7.15 – 8.00	6.99 – 8.18
$\sin^2 \theta_{12}/10^{-1}$ (NH or IH)	3.08	2.91 – 3.25	2.75 – 3.42	2.59 – 3.59
$\Delta m^2/10^{-3} \text{ eV}^2$ (NH)	2.44	2.38 – 2.52	2.30 – 2.59	2.22 – 2.66
$\Delta m^2/10^{-3} \text{ eV}^2$ (IH)	2.40	2.33 – 2.47	2.25 – 2.54	2.17 – 2.61
$\sin^2 \theta_{13}/10^{-2}$ (NH)	2.34	2.16 – 2.56	1.97 – 2.76	1.77 – 2.97
$\sin^2 \theta_{13}/10^{-2}$ (IH)	2.39	2.18 – 2.60	1.98 – 2.80	1.78 – 3.00
$\sin^2 \theta_{23}/10^{-1}$ (NH)	4.25	3.98 – 4.54	3.76 – 5.06	3.57 – 6.41
$\sin^2 \theta_{23}/10^{-1}$ (IH)	4.37	4.08 – 4.96 \oplus 5.31 – 6.10	3.84 – 6.37	3.63 – 6.59
δ/π (NH)	1.39	1.12 – 1.72	0.00 – 0.11 \oplus 0.88 – 2.00	—
δ/π (IH)	1.35	0.96 – 1.59	0.00 – 0.04 \oplus 0.65 – 2.00	—

